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Equation of State of Al-bearing Perovskite to Lower Mantle Pressure Conditions

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Abstract. Al_2O_3 is estimated to total 4 to 5 mol% in all mantle compositional models, and is believed to be incorporated into $(\text{Mg,Fe})\text{SiO}_3$ -perovskite at lower mantle conditions. Using synchrotron X-ray diffraction, we have measured the 300 K equation of state of a perovskite with $X_{\text{Al}} = \text{Al}/(\text{Al}+\text{Mg}+\text{Si}) = .077$, up to 32 GPa. A least squares refinement of two independent data sets yields the following equation of state parameters $V_0 = 163.52(3) \text{ \AA}^3$, $K_0 = 229(4) \text{ GPa}$, and $K'_0 = 2.5(4)$. These values allow us to assess the most probable substitution mechanism for aluminium into perovskite at pressure conditions of the shallow lower mantle. Al^{3+} is likely to be incorporated in place of Si^{4+} in the octahedral site of perovskite, hence requiring the creation of oxygen vacancies for charge balance. As a consequence, aluminous perovskites may have a high affinity for water. The enhanced compressibility of aluminous perovskite certainly has also strong geophysical and geochemical implications, as it may revive perovskite-rich lower mantle models.

Introduction

Much effort has been devoted in the last five years to characterize the thermal equation of state of the main phases constitutive of the lower mantle up at pressures and temperatures prevailing at these depths. Within the frame of a three components system $(\text{MgO}-\text{FeO}-\text{SiO}_2)$, all recent experimental work or calculations on $(\text{Mg,Fe})\text{SiO}_3$ -perovskite and $(\text{Mg,Fe})\text{O}$ magnesiowustite yield convergent results that a pure perovskite lower mantle is either very unlikely or very hot [Gillet *et al.*, 2000]. At temperatures most commonly accepted for the lower mantle, silicate perovskite would have bulk modulus values significantly higher than those predicted from the bulk modulus K_T profile of PREM [Dziewonsky and Anderson, 1981]. When minor components are further taken into account in calculating the thermoelastic profile of a lower mantle mineralogical model, Ca is often the next one considered [Wang *et al.*, 1996; Shim *et al.*, 2000]. Since Al does not enter a separate aluminous phase in the lower mantle, but is rather dissolved into $(\text{Mg,Fe})\text{SiO}_3$ -perovskite, with lesser amounts accepted into CaSiO_3 -perovskite [Irfune, 1994], Al has always been ignored in such models.

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However, various experimental data have indicated that the presence of Al in perovskite causes significant changes in the chemical and physical properties of silicate perovskite. Due to Al, the stability field of perovskite is extended toward iron-rich compositions [Kesson *et al.*, 1995]. Al-bearing perovskite has essentially the same $\text{Fe}/(\text{Fe}+\text{Mg})$ ratio as coexisting magnesiowustite, whereas Fe has a stronger affinity for magnesiowustite in the Al-free system [Wood and Rubie, 1996]. Al has also been shown to enhance the proportion of ferric iron in perovskite [McCammon, 1997], yielding an electrical conductivity about 3.5 times greater than that of an Al-free perovskite [Xu *et al.*, 1998].

Finally, Zhang and Weidner [1999] recently measured the effect of Al on the equation of state of perovskite. For an Al-content typical of the pyrolite composition, $X_{\text{Al}} = \text{Al}/(\text{Mg}+\text{Si}+\text{Al}) = .05$, they showed that the bulk modulus $K_{0,T}$ of Al-perovskite is about 10% smaller than that of the end-member MgSiO_3 -perovskite, whereas its temperature derivative is much larger. Since such results change dramatically the constraints on the mineralogical composition of the lower mantle, we present here the equation of state to lower mantle pressure conditions, for a slightly more aluminous magnesium silicate perovskite characterised by $X_{\text{Al}} = .077$, which will be designated as $\text{Al}_{.077}\text{-pv}$.

Experimental Procedure

Angle dispersive X-ray diffraction experiments were carried out in a membrane-type diamond-anvil cell at the ID30 high-pressure beamline of the ESRF (Grenoble, France). The incident monochromatic X-ray beam ($\lambda = 0.3738 \text{ \AA}$) was doubly focused and filtered to achieve a $10 \times 15 \text{ \mu m}^2$ spot on the sample. The diffracted signal was collected for 20 to 120 s, over a 2θ interval from 4 to 25° either on a MAR345 or on the FastScan detector.

Two P-V data sets were acquired for $\text{Al}_{.077}\text{-pv}$ samples synthesized from a $\text{Py}_{31}\text{Mj}_{69}$ synthetic glass containing 7.7 mol% Al_2O_3 . The starting glass was transformed into perovskite at 35 GPa and $1800 \pm 200 \text{ }^\circ\text{C}$ by laser heating with a CO_2 laser. The first sample was synthesized using potassium bromide as a pressure transmitting medium; it was afterward unloaded and triply washed in distilled water. The recovered 15 \mu m thick perovskite platelet was subsequently reloaded with ruby chips and He as a pressure transmitting medium. Diffraction patterns were then collected upon raising pressure to 24 GPa. The second sample was loaded

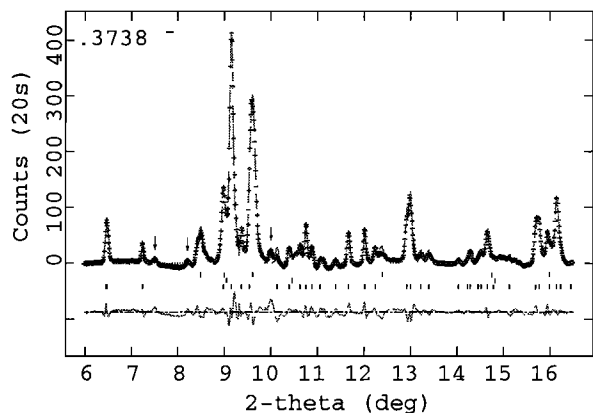


Figure 1. Integrated angle dispersive diffraction pattern of $\text{Al}_{0.77}\text{-pv}$ collected at 32 GPa. The difference between experimental intensities and the best LeBail refinement profile is shown at the bottom. Three series of ticks corresponding to the reflection lines of perovskite (bottom), fcc-argon (middle) and hcp-argon (top) are represented. Arrows point to weak unrefined reflections which could correspond to small amounts of stishovite and ilmenite (less than 1%).

cryogenically in Ar, with ruby chips. The glass was transformed into perovskite at 35 GPa; diffraction patterns were then recorded along the decompression path.

Two-dimensional images were integrated after spatial distortion and background corrections using the program Fit2D [Hammersley, 1997], and processed using the GSAS package [Larson and Von Dreele, 1988]. Le Bail full-profile refinements were applied to the diffraction patterns, in order to obtain reliable high-pressure cell parameters for Al-bearing perovskite (Fig. 1).

Results and Discussion

Axial Compressibilities and Distorsion

The variation of unit cell parameters as a function of pressure¹ is plotted in Fig. 2. At ambient pressure, $\text{Al}_{0.77}\text{-pv}$ lattice parameters are consistent with the dependence on Al content previously reported for perovskite [O'Neill and Jeanloz, 1994; Irifune et al., 1996; Kubo and Akaogi, 2000]. Substitution of 7.7% Al expands the structure by $\sim 0.7\%$, and causes an increased distortion in both a/b and a/c ratios relative to MgSiO_3 . As pressure is raised, the present data show that $\text{Al}_{0.77}\text{-pv}$ is elastically strongly anisotropic, with the b axis being approximately 30% less compressible than a and c (Fig. 2). Since all unit-cell parameters do not decrease linearly between 0 and 32 GPa, we calculate the axial compressibilities between 0 and 10 GPa, which allows comparison with the available data (Table 1). Whereas the b axis compressibility seems almost insensitive to Al substitution, we do observe a strong linear increase of both a and c compressibilities with Al content.

Reported in Fig. 3 are the variations of cell edge ratios a/b and $\sqrt{2}a/c$ as functions of pressure. These ratios may be related to the commonly used θ and Φ angles, needed

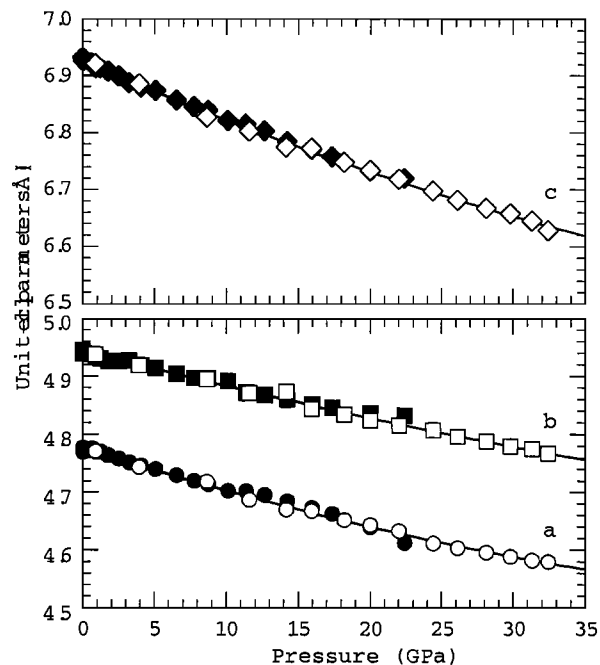


Figure 2. Pressure dependence of the unit-cell parameters of $\text{Al}_{0.77}\text{-pv}$. Circles, squares and diamonds represent respectively the a , b and c cell-edges. Filled and empty symbols designate data acquired respectively in He and Ar. The solid lines are guides-for-the-eye.

to describe the coupled octahedral rotations about the respective [110] and [001] pseudocubic axes. Assuming a rigid behaviour for the octahedra, departure of these rotational angles from 0 provides an approximate representation of the orthorhombic distortion in perovskite. At ambient condition, $\text{Al}_{0.77}\text{-pv}$ structure is characterised by $\theta = 14.9^\circ$ and $\Phi = 12.7^\circ$, and is thus slightly more distorted than MgSiO_3 perovskite. With increasing pressure, θ first increases rapidly up to ~ 10 GPa, and then rises slower than in MgSiO_3 perovskite, leading to an overall linear behaviour of $\cos\theta$ similar to that of MgSiO_3 perovskite over the entire investigated pressure range. At difference, Φ remains almost constant in $\text{Al}_{0.77}\text{-pv}$ between 0 and 32 GPa, while it slightly increases in Al-free perovskite. This indicates a common compression behaviour in the basal plane for Al-free and Al-bearing perovskites, but a distinct compression mechanism along the c axis.

Table 1. Linear compressibilities, bulk modulus, and their dependence on the Al content of perovskite in the 0-10 GPa range.

X_{Al}	β_a 10^{-3} GPa^{-1}	β_b 10^{-3} GPa^{-1}	β_c 10^{-3} GPa^{-1}	K_0 GPa
0 [†]	1.30(5)	1.04(4)	1.24(4)	254(13)
.05 [†]	1.39(4)	1.20(4)	1.42(5)	234(3)
.077	1.49(3)	1.02(4)	1.54(3)	224(3)
d/dX_{Al}	2.4(5)	0.0(3)	3.9(2)	-3.9(1) 10^2

[†]from Ross and Hazen [1990]

[‡]from Zhang and Weidner [1999]

¹Supporting material is available via Web browser or via Anonymous FTP from ftp://kosmos.agu.org, directory "apend" (Username = "anonymous", Password = "guest"); subdirectories in the ftp site are arranged by paper number. Information on searching and submitting electronic supplements is found at http://www.agu.org/pubs/esupp/about.html.

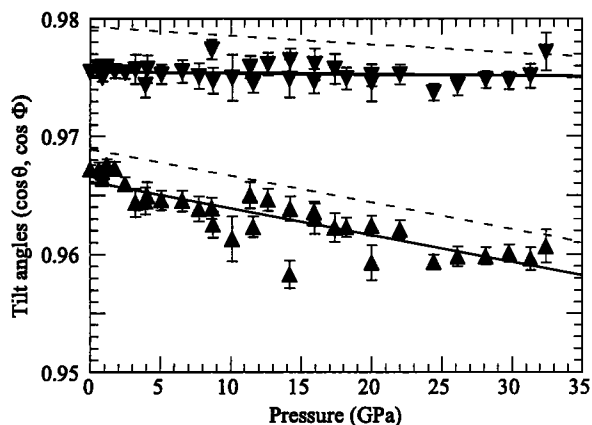


Figure 3. Variation of $\cos \theta = a/b$ (up triangles) and $\cos \Phi = \sqrt{2}a/c$ (down triangles) for $\text{Al}_{0.077}\text{-pv}$ as a function of pressure. A linear fit to the data (solid line) indicates that Φ remains almost constant up to 32 GPa. The evolution of the tilt angles for MgSiO_3 -perovskite (dashed line) is reported for comparison.

Equation of State

The variation of the unit-cell volume of $\text{Al}_{0.077}\text{-pv}$ with pressure is shown in Fig. 4. The pressure-volume data from the two runs were combined for a least square fit to a Birch-Murnaghan equation of state truncated to the 3rd order. Using all of the data, a bulk modulus $K_0 = 229(4)$ GPa, a pressure derivative $K'_0 = 2.5(4)$ GPa and a room pressure unit-cell volume $V_0 = 163.52(3)$ Å³ are obtained. Both values of K_0 and K'_0 for $\text{Al}_{0.077}\text{-pv}$ are much lower than those characteristics of MgSiO_3 perovskite [Fiquet *et al.*, 2000]. This demonstrates that Al-bearing perovskite is much more compressible than Al-free perovskite, at least up to pressures characteristic of the shallow lower mantle (Fig. 4).

Since all bulk moduli available for Al-bearing perovskites have been determined assuming K'_0 equal to 4 [Zhang and Weidner, 1999; Brodholt, 2000], we also adjusted our data to a 2nd order Birch-Murnaghan equation between 0 and 10 GPa. We calculate $K_0 = 224(3)$ GPa and $V_0 = 163.52(3)$ Å³, when K'_0 is set to 4. This value of the bulk modulus is in excellent agreement with that obtained by Zhang and Weidner [1999], and therefore indicates a steep decrease of the bulk modulus as the aluminium content in perovskite

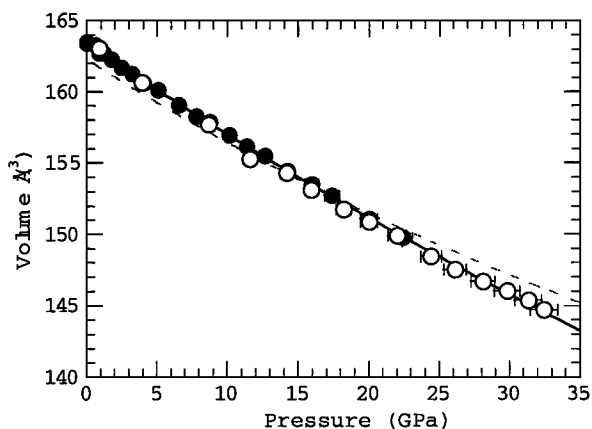


Figure 4. Birch-Murnaghan equation of state of $\text{Al}_{0.077}\text{-pv}$ (solid line). Filled circles: experiment in He; empty circles: experiment in Ar. $\text{Al}_{0.077}\text{-pv}$ becomes denser than pure MgSiO_3 -perovskite (dashed line) above approximately 17 GPa.

increases (Fig. 5). We derive a slope of $-3.9(1) \cdot 10^2$ GPa for the dependence of the bulk modulus of perovskite on the aluminium fraction X_{Al} (Table 2).

The large difference between both elastic parameters K_0 and K'_0 for pure MgSiO_3 -perovskite and $\text{Al}_{0.077}\text{-pv}$, together with their distinct anisotropy and pressure-induced crystallographic distortion tend to indicate that the compression mechanism of the perovskite structure evolves significantly even when a small amount of Al is dissolved. This suggests that the crystal chemistry of aluminous perovskites might be rather different from MgSiO_3 -perovskite.

Implications for the Substitution Mechanism of Al into Perovskite

Two substitution mechanisms have been proposed for the incorporation of Al^{3+} into perovskite. The first is the coupled substitution $2\text{Al}^{3+} \rightarrow \text{Al}_{\text{Mg}} + \text{Al}_{\text{Si}}$ where Al enters both cation sites. Through the second mechanism, Al replaces Si in the octahedral site only, and oxygen vacancies are thus required for charge balance.

In order to test whether oxygen vacancies exist in aluminous perovskite, Brodholt [2000] recently calculated the compressibility of end-member defect-bearing perovskites using *ab initio* methods. For a perovskite with one Mg and one Si in the unit cell substituted with Al, $(\text{Mg}_3\text{Al})(\text{Si}_3\text{Al})\text{O}_{12}$, the structure does not require the creation of vacancies for charge balance, and the calculated bulk modulus is only 6% lower than that of Al-free perovskite. At difference, the oxygen-deficient perovskite brownmillerite $\text{Mg}_2\text{Al}_2\text{O}_5$, which has all Al^{3+} in the octahedral site instead of Si^{4+} , is much more compressible with a calculated K_0 of 155 GPa. The large dependence of the bulk modulus of perovskite on the aluminium content described in the above paragraph

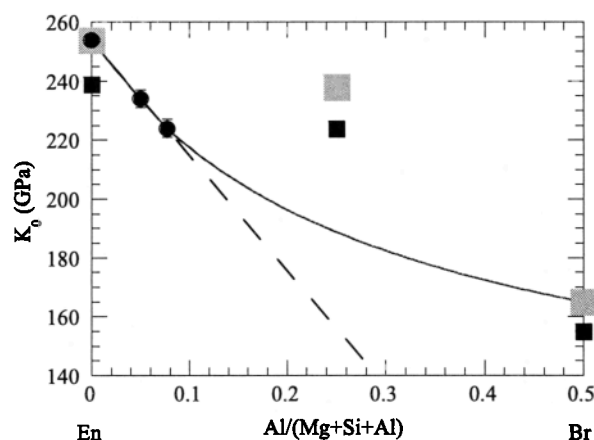


Figure 5. Evolution of the compressibility of perovskite as a function of its Al-content. All bulk moduli reported are valid within the 0-10 GPa range ($K'_0 = 4$). Filled circles include the bulk moduli experimentally determined by Ross and Hazen [1990], Zhang and Weidner [1999], and the present results. Empty squares depict the bulk moduli calculated by Brodholt [2000], for the vacancy-free MgSiO_3 (En), $(\text{Mg}_3\text{Al})(\text{Si}_3\text{Al})\text{O}_{12}$ perovskites, and for the oxygen deficient $\text{Mg}_2\text{Al}_2\text{O}_5$ brownmillerite (Br); solid grey squares represent the corresponding expected experimental values, typically 7-8% higher [Brodholt, 2000]. A linear fit to the experimental values (dashed line) indicates a very steep dependence of K_0 on the Al-content in perovskite. The solid line is an interpolation between available experimental values, and the only plausible value estimated at higher Al-content for brownmillerite.

provides the experimental confirmation of the hypothesis submitted by Brodholt [2000] that Al_2O_3 was likely to be introduced into MgSiO_3 perovskite in the form of the oxygen vacancy brownmillerite structure, rather than in place of Mg and Si through a coupled substitution (Fig. 5).

Geophysical implications

From the present results, we may conclude that for alumina content characteristic of the lower mantle (~5 mol%), all aluminium could be incorporated in perovskite by an oxygen vacancy mechanism, at least in the shallower part of the lower mantle [Brodholt, 2000]. The presence of Al^{3+} in the octahedral site and of oxygen vacancies in perovskite has various implications for both the geochemistry and the geophysics of the top of the lower mantle.

Aluminous perovskite might incorporate significant amount of hydroxyl, if as in ceramic proton-conducting perovskites, hydration of vacant oxygen sites is energetically favorable [Navrotsky, 1999]. This would strongly contrast with pure MgSiO_3 perovskite which does not dissolve any detectable amount of water [Bolfan-Casanova et al., 2000]. Hence, aluminous perovskite might absorb water from the subducting slabs which penetrate the 660 km discontinuity [Van der Hilst, 1995], and the upper part of the lower mantle might be hydrated.

Finally, the chemical compositions of the lower mantle deduced from mineral physics considerations have been shown to be extremely sensitive to thermoelastic properties of the perovskite phase. The lower mantle has been shown to be either consistent with a pyrolytic composition [Wang et al., 1994; Fiquet et al., 1998], or with a distinct composition enriched in Si [Stixrude et al., 1992]. Some consensus exists that, due to the high incompressibility of silicate perovskite compared to that of the lower mantle, only quite hot lower mantle models would be consistent with perovskite-enriched compositions [Gillet et al., 2000]. Smaller bulk modulus values induced by the presence of aluminium in silicate perovskite could thus revive the pertinency of silica-rich lower mantle models and of silicate earth stratification as far as chemical composition in major elements is concerned [Javoy, 1995]. In any case, the present results indicate that it is now absolutely necessary to take into account the elastic effects of Al in perovskite for assessing the Si/Mg ratio of the lower mantle, and thus the large-scale stratification of the Earth's mantle.

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