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Validation of Aura Microwave Limb Sounder BrO observations in the stratosphere

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[1] Validation of stratospheric BrO vertical profiles obtained by the Microwave Limb Sounder (MLS) on the Aura satellite is discussed. MLS BrO measurements are compared with expectations of its latitudinal and seasonal dependence, as well as with more localized balloon-borne measurements of BrO. We describe the expected precision and systematic errors of the version 2.2 retrieval and show that scientific studies using MLS BrO vertical profiles require extensive averaging to increase the signal-to-noise ratio to useful values. A monthly zonal mean over a 10° latitude bin (about 3,000 individual profiles) results in a precision of approximately ± 4 ppt ($\sim 25\%$ of a typical daytime signal). Moreover, it is necessary to take day/night differences to remove large biases. The pressure range over which the data are considered useful is 10 to 3.2 hPa. Over this range, the estimated accuracy in the day/night difference is about $\pm 20\%$. The vertical resolution is 5.5 km for 10 to 3.2 hPa. Day/night differences are a good measure of daytime BrO from 10 to 4.6 hPa; for 3.2 hPa the nonnegligible nighttime BrO needs to be accounted for. We infer total inorganic bromine (Br_y) to be 22.1 ± 5.5 ppt on the basis of analysis of MLS measurements of BrO, which implies a contribution of 6.5 ± 5.5 ppt to stratospheric bromine from sources other than long-lived CH_3Br and halons.

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1. Introduction

[2] Bromine monoxide (BrO) is an important species in the destruction of stratospheric ozone, especially in the midlatitude lower stratosphere [e.g., Yung *et al.*, 1980; Salawitch *et al.*, 2005; Sinnhuber *et al.*, 2006; World Meteorological Organization (WMO), 2007]. Model studies estimate that catalytic cycles involving BrO account for up to half of the chemical loss of polar ozone [e.g., McElroy *et*

al., 1986; Frieler *et al.*, 2006; WMO, 2007]. Yet there remains considerable uncertainty in the atmospheric abundance of BrO [WMO, 2007]. Bromine monoxide is the main daytime constituent ($\sim 50\%$) of stratospheric inorganic bromine (Br_y). Since relatively few measurements have been made of the other constituents, Br_y is typically estimated from stratospheric measurements of BrO, combined with estimates of the other constituents from photochemical models (often constrained by simultaneous measurements of NO_2 and O_3 [e.g., Sioris *et al.*, 2006]). One current problem is that estimates of Br_y inferred from measurements of stratospheric BrO obtained by various satellite and balloon-borne instruments disagree, ranging from 18 to 25 ppt [WMO, 2007, chap. 2]. Moreover, estimates of Br_y from measurements of stratospheric BrO exceed the amount (16–17 ppt) that can be provided by long-lived organic source gases (CH_3Br and halons). It is currently thought that very short lived substances (VSLS), many of biogenic origin, are an additional source of stratospheric inorganic bromine that can reconcile the bromine budget [WMO, 2007, chap. 2].

[3] Here we present a validation of stratospheric measurements of BrO by the Microwave Limb Sounder (MLS) instrument [Waters *et al.*, 2006] on the Aura satellite. This instrument has been measuring vertical profiles of BrO with

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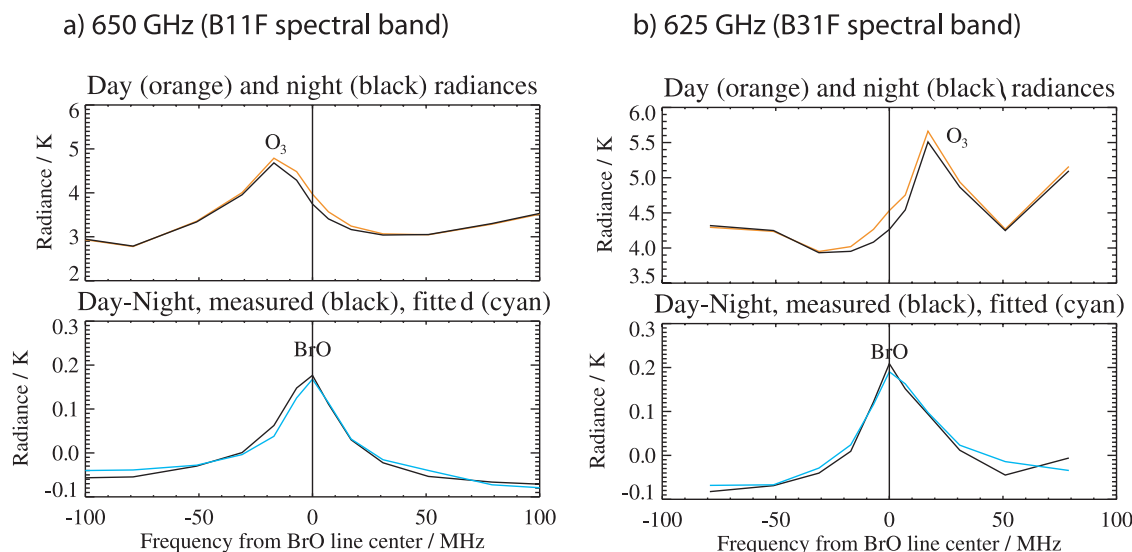


Figure 1. (a) Radiance spectrum of atmospheric emission in the region of the 650 GHz BrO line, MLS spectral band B11F, as detected by the R4 (640 GHz) radiometer, at an altitude of 40 km. The spectrum is averaged over 1 year, over latitudes ranging from 55°S to 55°N, and from 10 to 2.2 hPa. (b) Same as in Figure 1a except for the 625 GHz BrO line, MLS spectral band B31M. (top) Both the daytime (orange) and nighttime (black) radiances, revealing an overlapping ozone line. (bottom) Difference between the day and night signals (black), revealing the BrO line. The cyan line is the spectrum predicted from the measured profiles.

global coverage since August 2004. The latest publicly released version 2.2 data are analyzed here. We discuss the precision and accuracy of these data, examine how reasonable they are by comparing with expectations of dependences on latitude and season, and present preliminary comparisons with balloon-based measurements. Because the abundance of BrO varies strongly with solar zenith angle and with ambient levels of ozone and nitrogen dioxide, we also compare total inorganic bromine (Br_y), which is inferred from the BrO measurements using a photochemical model.

[4] The Aura MLS BrO data have already been used in a scientific study [Livesey *et al.*, 2006a]. However, that study did not use the version of data available publicly at that time, v1.5, because v1.5 BrO was determined to be unsuitable for scientific use. Rather, an interim retrieval version was used, in which a better choice was made in the tradeoff between precision and vertical resolution. Here we evaluate the latest publicly available version of BrO retrievals, v2.2, and compare them with the previous version, v1.5, as well as with the interim version used by Livesey *et al.* [2006a].

2. Description of Aura MLS Observations of BrO

[5] The Aura satellite orbits the Earth about 14 times per day. The MLS instrument detects thermal emission from the Earth's limb in the millimeter- and submillimeter-wavelength ranges as the field of view is scanned vertically through the atmosphere in the plane of satellite motion. Each vertical scan takes about 24 s and consists of 120 different forward looking limb radiance measurements, each measurement having an integration time of about 1/6th of a second. About 3,500 vertical scans are performed each day. At a given latitude, about half of these measurements are of a sunlit atmosphere (during the ascending part of the orbit) and half

are of a dark atmosphere (the descending part), except in regions of polar summer or winter, in which case both the ascending and descending measurements are in daylight (polar summer) or darkness (polar winter). The local solar time of the MLS measurements is about 1:45 p.m. for the ascending part of the orbit, and 1:45 a.m. for the descending part, except poleward of about 60° in latitude. The range of latitudes sampled is from 82°S to 82°N.

[6] BrO is measured using spectral lines of the ⁸¹BrO isotope that correspond to two rotational transitions in the vibrational ($v = 0$) and electronic ($J = 3/2$) ground state, $N = 47/2 \rightarrow N' = 49/2$, and $N = 49/2 \rightarrow N' = 51/2$. These transitions, at 624.768 GHz and 650.179 GHz, respectively, are each split into eight lines by Λ -doubling and nuclear quadrupole effects, and are spread over a range of about 6 MHz. The spectral features are characterized by Drouin *et al.* [2001], the collision-broadened line widths are reported by Yamada *et al.* [2003].

[7] The MLS BrO signal strength is only about 0.2 K brightness temperature, well below the individual measurement noise of about 4 K. Thus significant averaging of individual measurements is required to obtain useful BrO signal-to-noise. Figure 1a (top) shows a typical spectrum of atmospheric emission in the region of the 650 GHz BrO line as detected by the MLS R4 (640 GHz) radiometer. This spectrum is averaged over 1 year of measurements, for latitudes ranging from 55°S to 55°N, at a limb/tangent altitude of about 40 km. Both daytime measurements (orange) and nighttime measurements (black) are shown. An interfering ozone line is clearly visible. Taking the difference between the daytime and nighttime signals, shown in Figure 1a (bottom), removes the ozone signal, which does not display a diurnal variation at this altitude. The signal from BrO, which has a strong diurnal variation at this altitude, is thus revealed. Comparison of the measured

Table 1. Meaning of Bits in the “Status” Field

Bit	Value ^a	Meaning
0	1	flag: do not use this profile (see bits 8–9 for details)
1	2	flag: this profile is “suspect” (see bits 4–6 for details)
2	4	unused
3	8	unused
4	16	information: this profile may have been affected by high altitude clouds
5	32	information: this profile may have been affected by low-altitude clouds
6	64	information: this profile did not use GEOS-5 temperature a priori data
7	128	unused
8	256	information: retrieval diverged or too few radiances available for retrieval
9	512	information: the task retrieving data for this profile crashed (typically a computer failure)

^a“Status” field in L2GP file is total of appropriate entries in this column.

spectrum (black) with that predicted from the retrieved profiles (cyan) shows good agreement. Figure 1b shows similar results for the other BrO line (at 625 GHz) that is detected by the same radiometer. Both lines are used for the BrO retrievals.

[8] The MLS retrieval technique has been described previously [Livesey *et al.*, 2006b]. Profiles are retrieved on a pressure grid of six levels per decade change in pressure (~ 2.5 km vertical spacing). Retrieved values of BrO are not constrained to be positive because to obtain a useful signal-to-noise ratio, extensive averaging is required; were the values constrained to be positive, the average would then be biased positive.

[9] Data are stored in HDF-EOS version 5 “swath” format, with each file containing data for a 24 h period from midnight to midnight universal time. The MLS Version 2.2 data quality document (a preliminary version available at <http://mls.jpl.nasa.gov/data>) gives more information on the format and content of these data files. Each MLS retrieved data value is stored along with a corresponding precision value that quantifies the impact of MLS radiance measurement noise on the data. As will be discussed below in the section on precision, the value of the precision is made negative if it is worse than 50% of the value of the a priori precision. Each profile is classified with three fields indicating the quality of the data. The “Quality” field indicates how well the measured radiances can be fit by the retrieval algorithm; higher values indicate a better fit. The “Status” field contains several flags indicating whether there were enough measured radiances, whether clouds might have affected the data, and whether there might be other reasons not to use the data. These Status flags are shown in Table 1. The “Convergence” field indicates how the radiance fit compares with that expected by the retrieval algorithm. Values near 1.0 indicate good agreement. We recommend only using v2.2 BrO data that meet the following criteria: (1) precision value for data point is positive, (2) status field for profile is even, (3) quality for profile is greater than 1.2, and (4) convergence for profile is less than 1.5.

[10] Version 2.2 (v2.2) is the 2nd public release of MLS data and has been used to process the incoming data stream since March 2007. Reprocessing of the data collected since MLS became operational in August 2004 is also in progress using the v2.2 algorithms. These processing streams have the specific version name v2.21 and include a minor software patch applied to an earlier version v2.20 that corrects the handling of MLS Level 1 radiances flagged

as bad data. We refer to both these versions collectively as version 2.2 (v2.2). For this validation effort, only a small subset of days covering late 2004 to early 2007, about 93, has been processed with v2.20, with priority given to days on which there are correlative measurements.

[11] Averaging kernels [Livesey and Read, 2000; Rodgers, 2000; Livesey *et al.*, 2006b] for the v2.2 BrO product are shown in Figure 2. They indicate the vertical resolution of the retrieved product for each pressure level, as well as the dependence of the retrieved product on the a priori. The dashed black line gives the width of each curve at half maximum, a measure of the vertical resolution, which for the pressure range of scientifically useful MLS BrO ranges from ~ 5.5 km at 10 hPa to ~ 6.0 km at 3.2 hPa. Values are given in Table 2. The solid black line in Figure 2 is the sum of the averaging kernel rows for each pressure surface and indicates the relative amount that the retrieval is based on the MLS measurement as opposed to being based on the a priori; a value near unity indicates that the retrievals are based almost entirely on the MLS measurement. For BrO, the profiles are effectively independent of the a priori for pressures of 0.32 hPa and greater.

3. Retrieval

[12] BrO vertical profiles from the first publicly available version, v1.5, are shown in Figure 3 (top). The daily zonal mean profiles (black) of the daytime and nighttime measurements (Figure 3, top left and top middle, respectively) display large oscillations with altitude, as well as strong biases at low altitude. Taking the average (red) of these profiles lessens the range of oscillation, while taking the day/night difference (Figure 3, top right) lessens the magnitude of the bias. A poor choice of tradeoff between precision and vertical resolution led to the oscillations, making v1.5 unsuitable for scientific studies. We then developed a simple interim retrieval version, shown by Figure 3 (middle), in which a better choice was made in the tradeoff between precision and vertical resolution. For this interim version, daily zonal mean radiances were calculated, from which vertical profiles of BrO were then retrieved [Livesey *et al.*, 2006a]. The most recent version, v2.2, is shown in Figure 3 (bottom). It differs from v1.5 in that smoother profiles are obtained at the expense of decreased vertical resolution. Moreover, new spectroscopy is used for the stronger ozone lines (located in a different spectral region than the BrO emission) (B. J. Drouin and R. R. Gamache, manuscript in preparation, 2007), which significantly reduces v1.5 biases that had been empirically

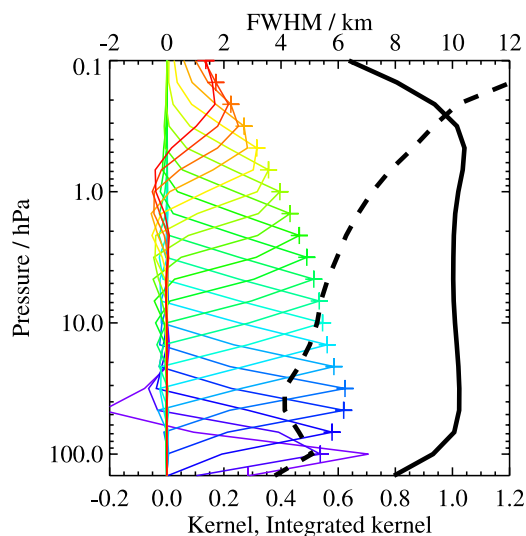


Figure 2. Vertical averaging kernels (integrated in the horizontal dimension for five along-track profiles) for MLS v2.2 BrO at a latitude of 35°N. Variation in the averaging kernels is sufficiently small that these are representative of all profiles. Each colored curve shows the averaging kernel for a particular MLS pressure surface (denoted by a plus sign in the corresponding color) and indicates the region of the atmosphere from which information is contributing to the measurements on that individual retrieval surface. The full-width at half maximum (FWHM) of each curve along the pressure axis is a measure of the vertical resolution of the retrieved profile for that pressure surface. That width, as measured in km (top axis), is shown by the dashed black curve. The solid black curve shows the sum of the averaging kernel columns for each of the pressure surfaces (i.e., is the integrated area under each two-dimensional kernel, both horizontal and vertical) and is an indication of how much of the retrieved product is based on the a priori. Values near unity imply that the majority of information for that MLS data point has come from the measurements, while values near zero imply the majority has come from the a priori.

ascribed to ozone. Compared to the old HITRAN spectroscopic constants, the new line widths are, on average, 8% more narrow at 296 K and show more variability in the temperature exponents. Changes in the line strengths and positions are negligible (less than a tenth of a percent). Figure 3 shows that for the pressure range over which MLS

measurements of BrO are useful, from 10 to 3.2 hPa (as will be justified below), the average nighttime v2.2 BrO (red) shows larger positive biases than does the average nighttime interim version BrO, while the day/night difference is comparable, with v2.2 showing about 3 ppt more BrO than the interim version. Table 3 summarizes the three retrieval versions.

4. Precision

[13] The expected precision in each retrieved profile is calculated from radiance noise and stored in the data files with each retrieval value. As mentioned earlier, as a data quality flag, the value of the precision is made negative if it is worse than 50% of the value of the a priori precision. Figure 4 compares the expected precision (thick line) of one measurement of BrO with the measured precision (circles) based on the scatter in the retrieved profiles, for both v1.5 and v2.2 retrievals. Also shown is the precision for daily, monthly, and yearly 10° zonal averages. The precision is seen to have improved in v2.2, at the expense of vertical resolution. As shown in Figure 4, a monthly zonal mean over a 10° latitude bin (about 3,000 individual profiles) results in an estimated precision of about ± 4 ppt (about 20% of a typical daily signal) over a pressure range from 10 to 3.2 hPa.

[14] Figure 5 shows a histogram of the distribution of daytime (shaded region) and nighttime (red line) BrO measurements for four different pressure surfaces over a latitude range from 55°S to 55°N for the days reprocessed in v2.2 to date. Also shown is a Gaussian fit (black line) to the daytime measurements, which are seen to have the Gaussian distribution expected from noise and a width that closely matches the estimated precision. The diurnal variation in BrO is discernible for the lower altitudes.

5. Accuracy

[15] We now address quantification of the various sources of systematic uncertainty in the MLS BrO product. Systematic errors (which are reproducible but cannot be lessened by averaging) can be introduced into the retrieved product by errors in instrument calibration (e.g., through radiometric calibration and field-of-view characterization), through spectroscopic uncertainties or approximations, and by approximations in the retrieval formulation and implementation. A comprehensive assessment of systematic errors and their effect on all the MLS products is given by *Read et*

Table 2. Summary of Aura MLS BrO Product^a

Region, hPa	Vertical Resolution, km	Precision, ^b ppt	Bias Uncertainty, ^c ppt	Scaling Uncertainty, ^c %	Comments
2.2 and less	–	–	–	–	unsuitable for scientific use
3.2	6	± 5	± 6	± 20	need to account for nonnegligible nighttime BrO
4.6	5.5	± 4	± 9	± 20	
6.8	5.5	± 4	± 20	± 20	
10	5.5	± 4	± 30	± 20	
150–15	–	–	–	–	unsuitable for scientific use
1000–215	–	–	–	–	not retrieved

^aBecause of large biases in the data, the daytime and nighttime BrO data at all pressure levels are unsuitable for scientific use. Rather, day/night differences must be used. Note that day/night differences are not useful for polar winter and summer, where BrO does not undergo a diurnal variation.

^bPrecision is for a 10° monthly zonal mean profile.

^cAccuracy is based on the systematic error tests.

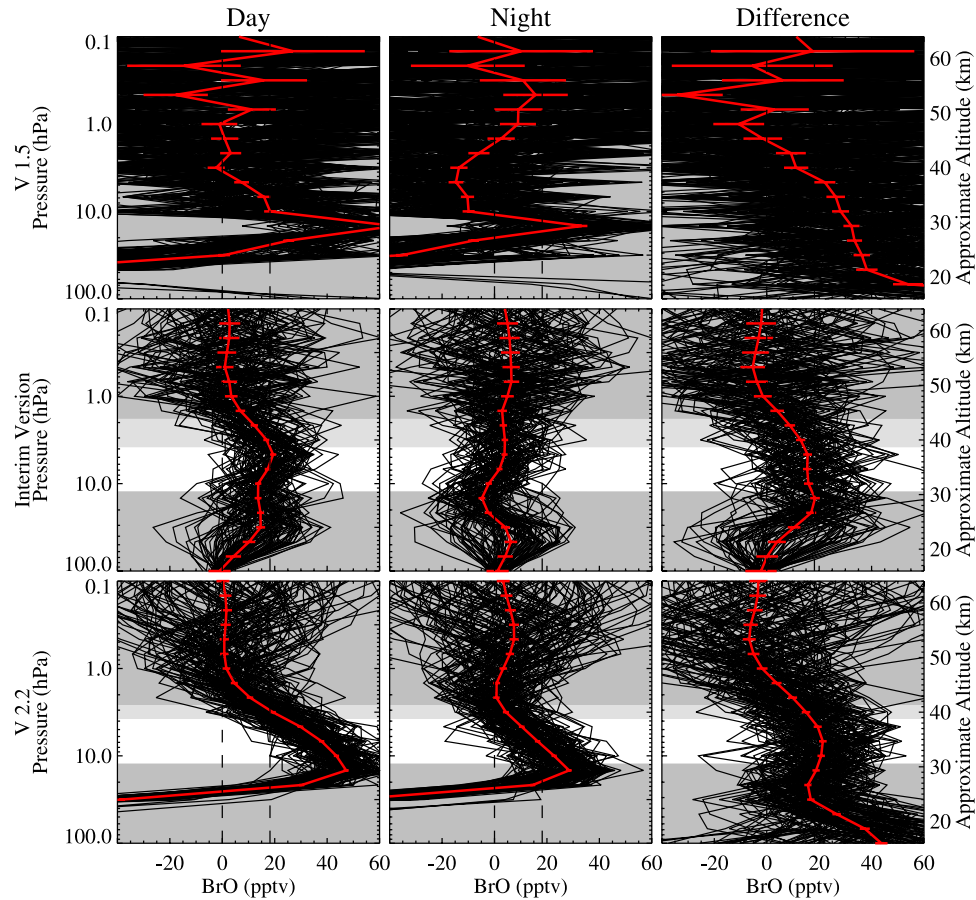


Figure 3. Vertical profiles for three retrieval approaches for daytime, nighttime, and day/night difference. Profiles are averaged over the subset of days (149) retrieved to date with v2.2 and over latitudes from 35°N to 45°N. (top) Version 1.5, (middle) the interim version, and (bottom) version 2.2. Black profiles are daily zonal means; red is the average. Error bars show precision only. (We note that the average for the interim version corresponds to the version 2.2 subset of retrieved days occurring before 5 March 2006; after this date, data for the interim version were not processed.) To make comparison easier, a dashed vertical line is drawn at 18.2 pptv, the value of the interim version average BrO at 10 hPa. Regions shaded dark gray indicate where MLS BrO has been determined to be unsuitable for scientific use. Regions shaded light gray indicate where nighttime BrO needs to be taken into account in the day/night difference for that difference to be a good measure of daytime BrO.

Table 3. MLS BrO Retrieval Versions

	1.5 ^a	Interim	2.2
Region, ^{b,c} hPa	10 to 1 ^a	10 to 2.2	10 to 3.2
Vertical resolution, km	2.5 ^a	6	5.5 to 6 ^d
Precision, ^{d,e} ppt	±20 ^a	±3	±4 to 5 ^d
Bias uncertainty, ppt	±3 to 10 ^a	–	±6 to 30 ^d
Scaling uncertainty, %	–	±30	±20
Comments	poor choice of tradeoff between precision and vertical resolution	zonal averages of radiances	better choice of tradeoff between precision and vertical resolution; new ozone spectroscopic constants

^aVersion 1.5 is unsuitable for scientific use at all pressures.

^bPressure range over which the day/night difference is scientifically useful. Because of large biases in the data, the daytime and nighttime BrO data at all pressure levels are unsuitable for scientific use; rather, day/night difference must be used. Note that day/night differences are not useful for polar summer and winter, where BrO does not undergo a diurnal variation.

^cFor pressures ≤ 3.2, the day/night difference is not a good measure of daytime BrO.

^dDepends on pressure range. See Table 2 for more details.

^ePrecision is for a 10° monthly zonal mean profile.

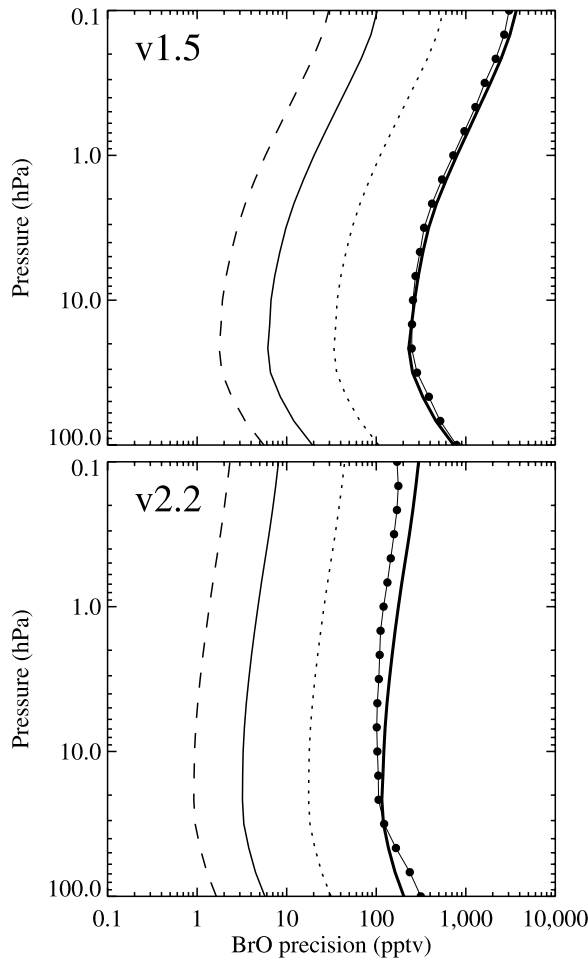


Figure 4. Comparison of the measured precision (circles) with that expected from the retrieval (thick solid line), for a single profile, for (top) v1.5 and (bottom) v2.2. The measured precision on a single profile was determined from the scatter in nighttime profiles retrieved over a single day within a 10° latitude bin. Also shown is the expected precision for the day/night difference of 10° zonal mean profiles averaged over a day (dotted line), a month (thin solid line) and a year (dashed line).

al. [2007, Appendix A]. In this section we summarize the results of this assessment relevant for the BrO product.

[16] The effect of each identified source of systematic error on MLS measurements of radiance has been quantified and modeled. These quantified effects correspond to either 2σ estimates of uncertainties in each MLS product, or an estimate of the maximum reasonable uncertainty based on instrument knowledge and/or design requirements. For each source of systematic error, we have used one of the following two methods to quantify the effect on the MLS retrieved products.

[17] In the first method, sets of modeled errors corresponding to the possible magnitude of each uncertainty have been applied to simulated MLS cloud-free radiances, based on a model atmosphere, for a whole day of MLS observations. These sets of perturbed radiances have then been run through the routine MLS data processing algorithms, and the differences between these runs and the

results of an “unperturbed” run have been used as a measure of the systematic uncertainty in each case. Although the term “systematic uncertainty” typically refers to an unknown additive bias or scaling factor, many of the sources of systematic uncertainty in the MLS measurements give rise to additional “scatter” in the products. For example, although an error in the ozone spectroscopy directly causes a bias on the ozone retrieval, it also has a secondary effect on the retrieval of other species with overlapping lines. The difference between the retrieved product in the unperturbed run and the original “truth” model atmosphere is taken as a measure of uncertainties due to retrieval formulation and numerics. The concentration in the unperturbed case is plotted against the concentration in the perturbed case. A linear regression produces an intercept and a slope. The intercept is a measure of errors that are independent of the concentration (additive bias), while the deviation of the slope from unity is a measure of errors that are proportional to the concentration (multiplicative error). When day/night differences are taken, the additive bias errors should cancel, while the multiplicative errors remain.

[18] In the second method, the potential impact of some remaining (typically small) systematic uncertainties has been quantified through calculations based on simplified models of the MLS measurement system [Read *et al.*, 2007]. These calculations provide estimates of possible multiplicative error.

[19] The first panel in Figure 6 shows an estimate of the potential magnitude of the additive biases introduced into the v2.2 BrO product by each of the sources of error listed in Figure 6. The first panel in Figure 7 zooms in on the pressure region over which MLS BrO measurements are useful. These values should be considered as 2σ estimates of their probable magnitude. The second panel in Figures 6

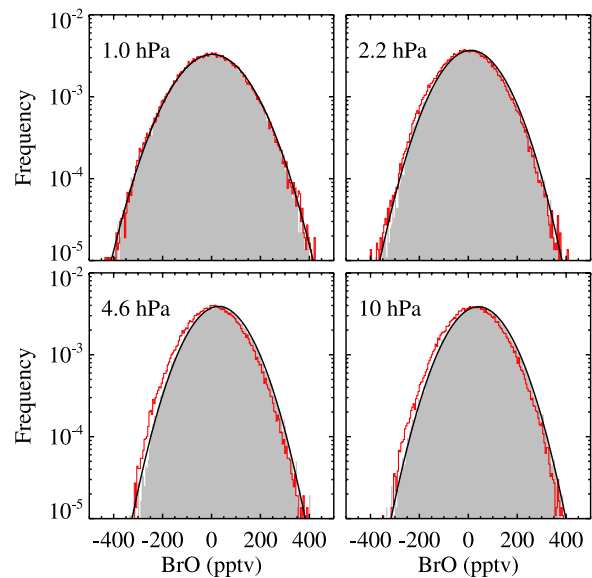


Figure 5. Histogram showing the distribution of daytime (shaded region) and nighttime (red curve) MLS BrO measurements at four different pressure surfaces for a latitude range from 55°S to 55°N for all the v2.2 data processed to date. The black line shows a Gaussian fit to the daytime data.

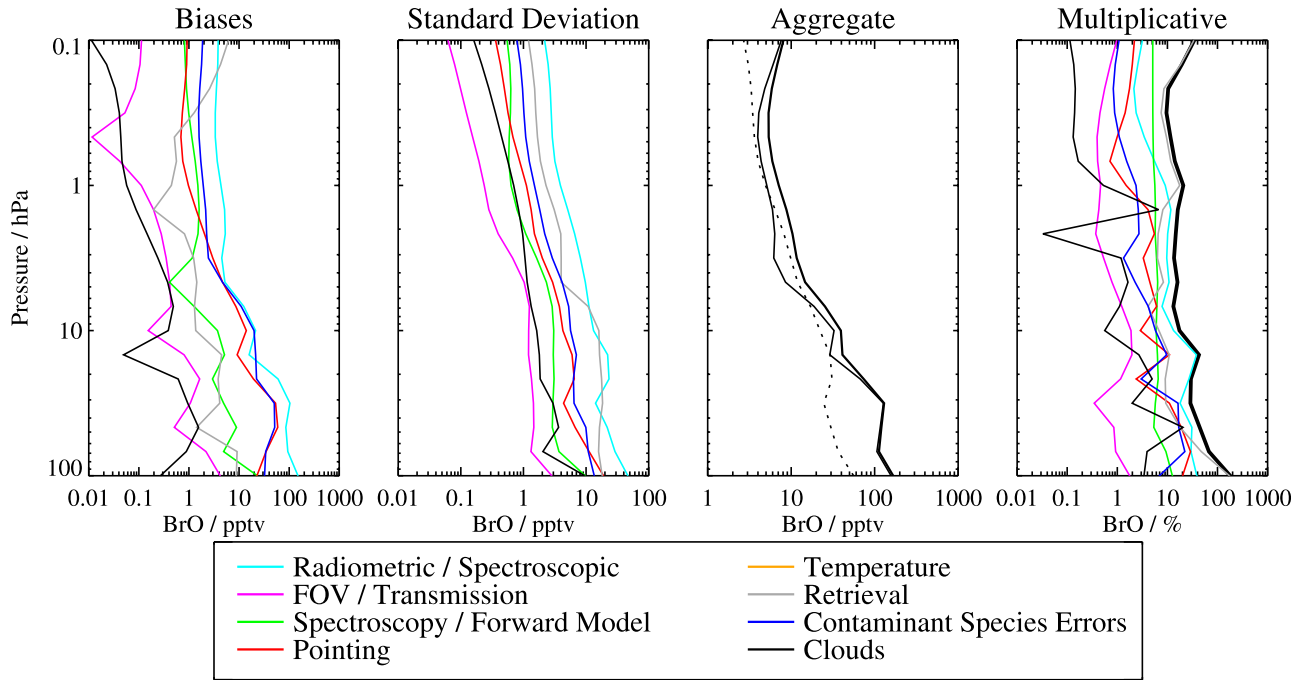


Figure 6. Estimated accuracy in v2.2 BrO measurements. The first panel shows the estimated potential magnitude of the systematic additive error, or bias, introduced into BrO by each of eight sources, while the second panel shows the additional scatter (see text) introduced into the retrieved values by each of the sources of uncertainty. The third panel shows the root sum squares (RSS) of all the biases (thin solid); the RSS of all the additional scatters, or standard deviations (dotted); and the RSS of the two (thick solid). The fourth panel shows the estimated potential multiplicative bias.

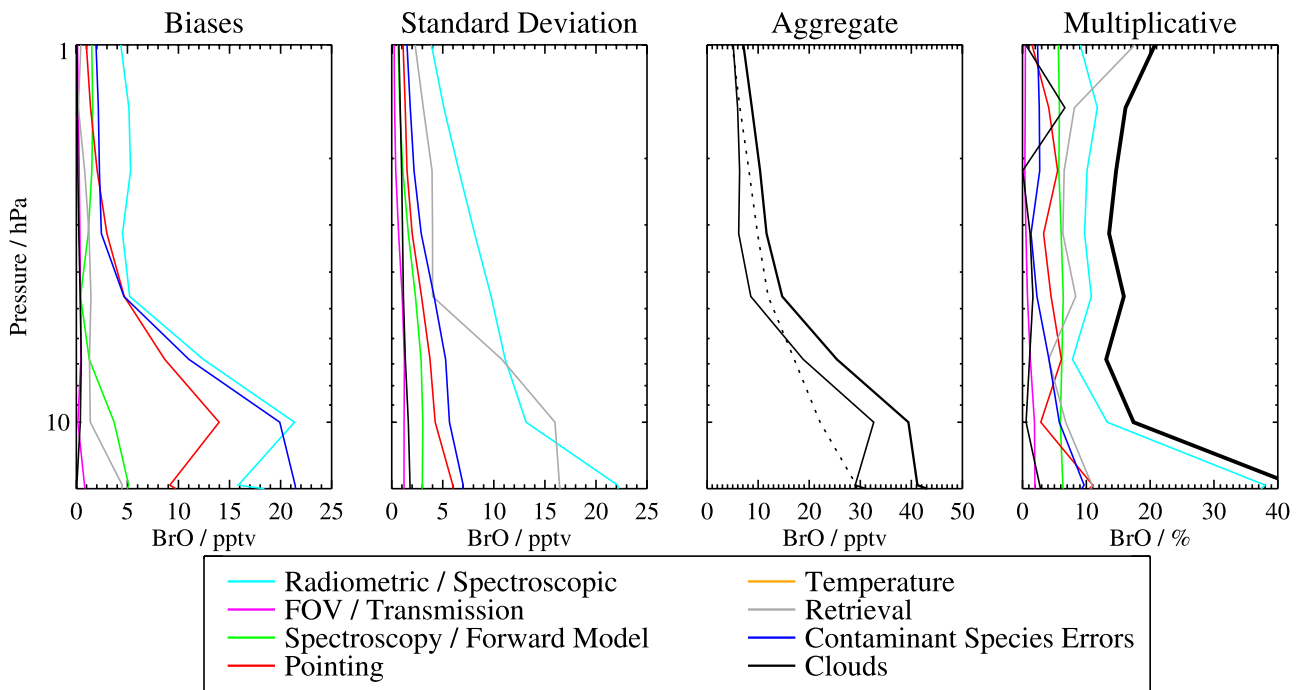


Figure 7. Same as Figure 6 except that it is a blowup of the pressure region over which MLS BrO is scientifically useful. Note that the x axis is linear.

and 7 shows additional scatter in the additive bias introduced by each of these sources.

[20] The dominant source of additive systematic error in the MLS BrO product is uncertainty in the MLS radiometric and spectroscopic calibration (cyan), and is mainly due to the spectral signature introduced in MLS radiances by departures from a linear response within the signal chain that are not accounted for in the calibration. This uncertainty alone can cause a potential additive bias in BrO of up to about ± 20 ppt at 10 hPa, and up to about ± 100 ppt at 31.6 hPa. A second major contributor to systematic errors in the BrO signal is uncertainty in overlapping signals from other species (blue), mainly ozone, as seen in Figure 1, though nitric acid also contributes. This uncertainty alone can cause a potential additive bias of up to ± 20 ppt at 10 hPa, and up to ± 50 ppt at 31.6 hPa. A third major contributor to additive bias is pointing (red), which is caused by the impact of errors in the assumed width of the oxygen line used in the MLS pointing retrieval, uncertainty in the field of view pointing offsets between the 640 GHz and 240 GHz radiometers, and the uncertainty in the field of view pointing offsets between the two 118 GHz radiometers and the 240 GHz radiometer. This uncertainty alone contributes up to about ± 15 ppt at 10 hPa, and up to ± 50 ppt at 31.6 hPa.

[21] The aggregate additive bias uncertainty in MLS BrO measurements, shown by the thin line in the third panel of Figures 6 and 7, can be as high as about ± 30 ppt ($\sim 400\%$ of the expected BrO signal) at 10 hPa, decreasing to about ± 6 ppt (50%) at 3.2 hPa. Because averages over large numbers of profiles are needed to obtain useful precision for MLS BrO, the additional scatter (dotted line) introduced into the data from these sources of uncertainty will become negligible, and thus the aggregate (thick line) of the biases and scatter will average down to just the biases (thin line).

[22] The fourth panels in Figures 6 and 7 show the multiplicative uncertainty (the accuracy as a percent of the signal, as opposed to an additive bias) introduced into the MLS BrO measurements from the sources of error listed above. For the pressure range over which MLS BrO measurements are useful, 10 to 3.2 hPa, this scaling uncertainty is up to about $\pm 20\%$ of the signal.

[23] We can dramatically reduce the effects of the additive bias in the BrO signal by taking advantage of the diurnal variation of BrO and subtracting the nighttime signal from the daytime signal. For pressures of 4.6 hPa and greater, our calculations show that the nighttime BrO is expected to be negligible, and therefore this difference is a measure of the daytime BrO. For lower pressures, the subtraction can still be done; however, for the difference to be a good measure of the daytime BrO, the nonnegligible nighttime BrO signal will need to be taken into account (perhaps with the aid of a photochemical model). While taking day/night differences minimizes the additive bias, which we treat as negligible, it does not affect the multiplicative uncertainty factor, which remains at up to about $\pm 20\%$.

[24] The uncertainty in the MLS BrO data due to systematic errors is summarized in Table 2. We conclude that it is necessary to take day/night differences in the measurements of BrO. Even so, because of the extremely large biases at 15 hPa and larger pressures, ($+30$ ppt at 15 hPa,

-40 ppt at 26 hPa, and becoming even more negative for higher pressures), we conclude that MLS BrO is not useful for scientific studies in this region. For pressures ranging from 10 to 4.6 hPa, and with sufficient averaging, the error analysis indicates that the MLS BrO abundance uncertainty is about $\pm 20\%$. For 3.2 hPa, the day/night difference can be known to within an uncertainty of 20%, but for this difference to be a good estimate of daytime BrO, one will need to compensate for the nonnegligible nighttime BrO. We note that this method of taking day/night differences is not applicable for polar summer and winter, where BrO is constant throughout the day and night.

6. Comparison With Expectations

[25] The SLIMCAT chemistry-transport model [Chipperfield, 1999] was run in “near-real time” and driven by U.K. Met Office analysis wind and temperature fields. Model results were sampled at the same locations and times as MLS profile observations. JPL 2002 kinetics [Sander *et al.*, 2003] were used, with the addition of the reaction $\text{BrONO}_2 + \text{O} \rightarrow \text{BrO} + \text{NO}_3$ [Soller *et al.*, 2001], which is included in the JPL 2006 kinetics compendium [Sander *et al.*, 2006]. The model was initialized with 16 ppt CH_3Br (long-lived source for inorganic bromine) plus 6 ppt of Br_y (in the form of BrO and BrONO_2) to represent the effect of short-lived sources at the 326 K model boundary; model results show that all of the bromine is in the form of inorganic bromine (Br_y) for pressures less than 30 hPa.

[26] Figure 8 shows seasonal zonal means of ascending (Figure 8a), descending (Figure 8b), and the day/night difference (Figure 8c) MLS v2.2 BrO as a function of pressure and latitude. The precision on the day/night differences depends on the number of measurement days in that season: about ± 6 ppt for SON 2004, DJF 2004/5, and MAM 2005; and about ± 10 ppt for JJA 2005. These plots can be compared with corresponding plots of SLIMCAT BrO, shown in Figure 9. The MLS BrO day/night difference shows the expected behavior with season and latitude. At midlatitudes and tropics, MLS BrO displays a diurnal variation, with more BrO during the day than at night. In the polar summer and winter regions, with constant daytime and nighttime, respectively, the MLS day/night difference is close to zero.

[27] We take a more quantitative look at the comparison of MLS BrO day/night differences and SLIMCAT values in the midlatitude and tropics for the four seasons by using a scatterplot, shown in Figure 10. As in Figure 9, the MLS data points represent seasonal zonal means averaged over 10° latitude bins. We see that although the MLS data points exhibit significant scatter, they are distributed about the SLIMCAT values. Implications for total inorganic bromine (Br_y) are discussed in section 8.

[28] Figure 11 shows vertical profiles of seasonal averages of MLS v2.2 BrO during polar summer in the northern hemisphere (NH) (Figure 11, top) and southern hemisphere (SH) (Figure 11, bottom), for ascending (daytime) measurements, descending (also daytime) measurements, and the ascending/descending difference. These are compared with SLIMCAT model profiles, similarly averaged. The large biases in the MLS BrO data (for both ascending and descending measurements) are evident. The origin of these

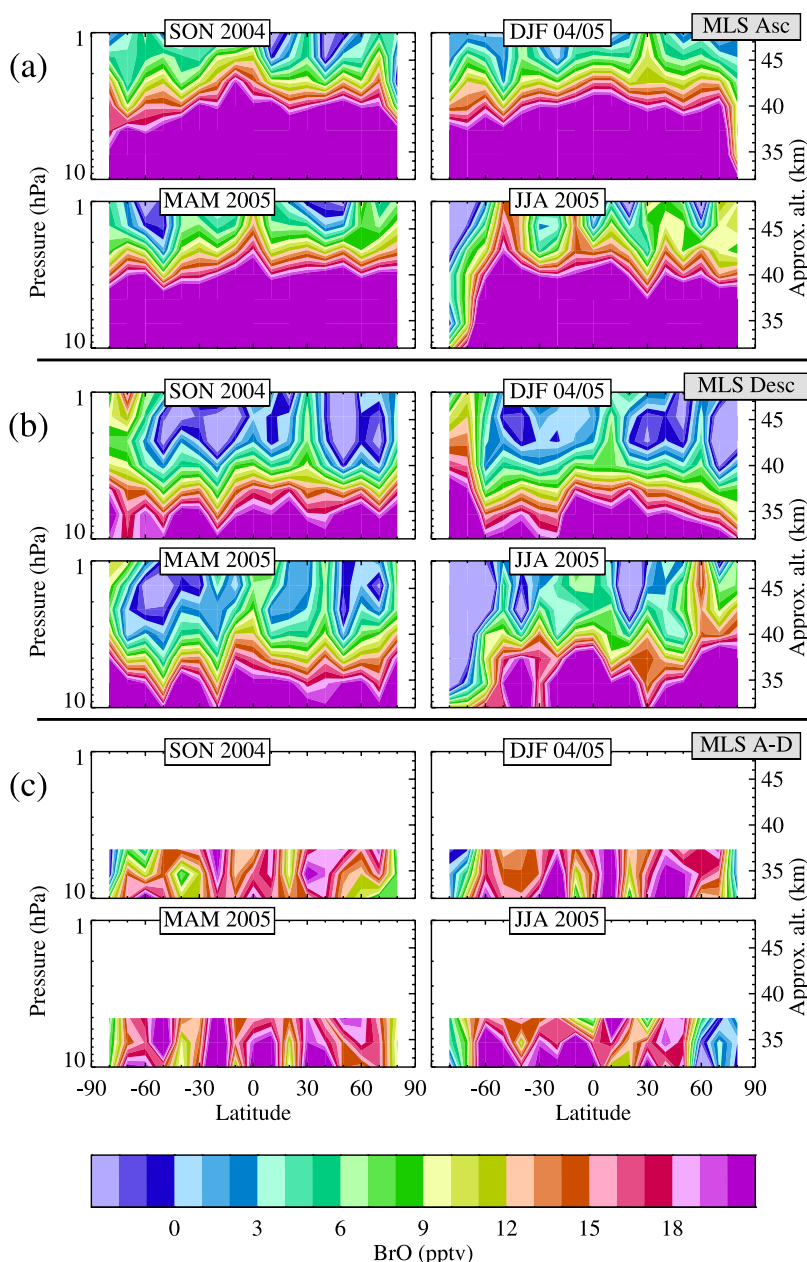


Figure 8. Seasonal zonal means of MLS BrO observations for (a) ascending (mainly daytime) and (b) descending (mainly nighttime) phases of the orbits. (c) To alleviate biases in the lower-altitude regions, the ascending/descending difference can be used as a measure of daytime BrO, so long as the expected nighttime abundance of BrO is negligible, which is the case for MLS difference data on the 4.6 hPa and greater pressure surfaces shown here.

is discussed in section 5 on accuracy. Over the pressure range 10 to 3.2 hPa, the MLS BrO difference profiles are zero within the expected uncertainty (i.e., taking the difference removes the biases apparent in the ascending and descending data). However, for lower pressures, a large deviation from zero is seen for the SH (Figure 11, bottom) difference profile, by as much as 10 ± 5 ppt at 1.5 hPa. Because this difference is not behaving as expected for these lower pressures, we define 3.2 hPa to be the lowest pressure for scientific use, though this may be overly cautious.

[29] Figure 12 shows the same as Figure 11 except for polar winter regions. For the NH polar winter (Figure 12, top), the SLIMCAT model results show that the difference is not expected to be zero, because over the days included in the seasonal average, there is some sunlight during the ascending part of the orbit. In this case, the MLS day/night difference also differs from zero.

[30] Figure 13 shows a time series of daily zonal mean MLS BrO over the whole mission for both v1.5 (black) and v2.2 (red) for descending (nighttime) measurements (Figure 13, top) and day/night differences (Figure 13, bottom). In Figure 13 (top), a discontinuity is seen in the

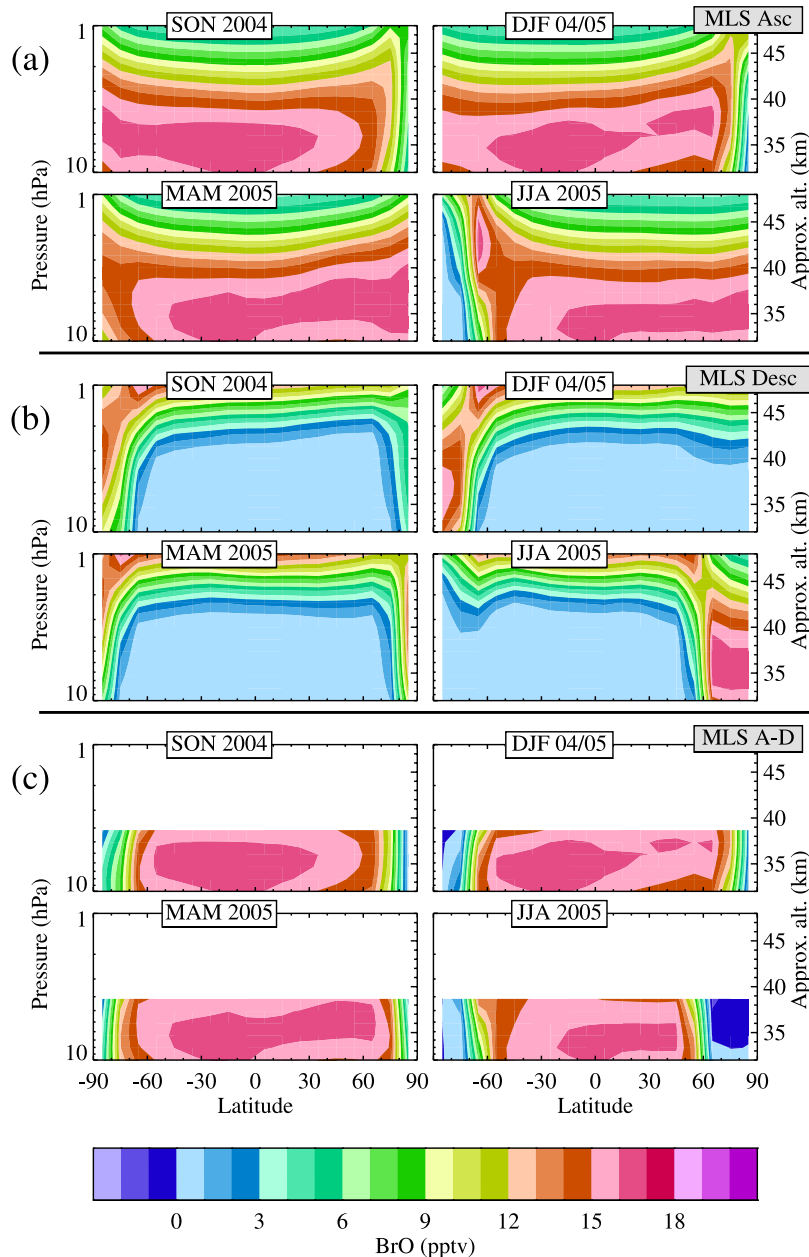


Figure 9. Same as Figure 8 except for SLIMCAT model estimates of BrO.

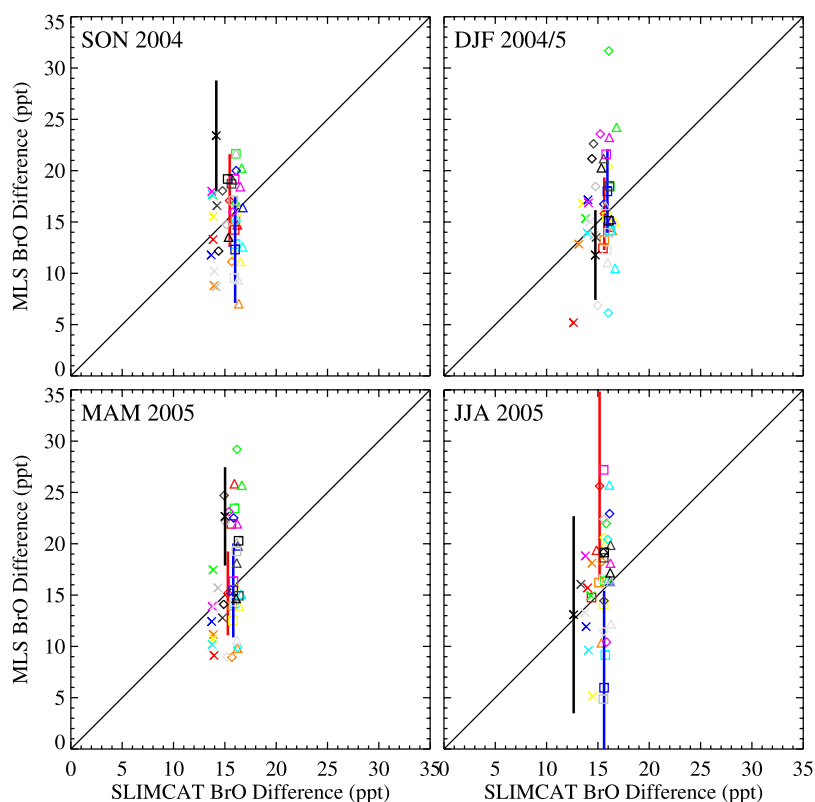
v1.5 descending (nighttime) data at about March 2006, which was caused by a change in the spectrometers when MLS band 13 [Waters *et al.*, 2006] was switched off to conserve measurement capability. While this band is not used to detect the BrO signal, it is used in the retrieval of other species and thus can affect the background signal. It can be seen that this discontinuity is removed by taking day/night differences, as shown in Figure 13 (bottom). As version 2.2 was designed subsequent to the band being switched off, it is reassuring to see that no corresponding discontinuity is seen in the v2.2 nighttime data.

7. Comparison With Other Measurements

[31] We now compare vertical profiles of MLS BrO with measurements by balloon-borne instruments. While better

comparisons would be obtained by taking monthly zonal means of MLS BrO profiles about the dates and latitudes of each flight, the required v2.2 data are not yet available. Thus, to achieve reasonable precision, we take an average of all the days (149) of MLS v2.2 reprocessed to date, over a zonal mean from 55°S to 55°N. This latitude range is chosen to ensure that measured BrO undergoes a diurnal variation. This average samples all seasons and a large range of solar zenith angles, 26° to 73°.

[32] Because BrO abundances vary strongly with solar zenith angle, as well as with local abundances of O₃ and NO₂, we also compare vertical profiles of total inorganic bromine (Br_y), which are not as sensitive to local conditions. Values of Br_y are inferred from the BrO measurements using a photochemical model [Osterman *et al.*, 1997]. Importantly, since all the organic bromine is



0

Figure 10. Scatterplot of midlatitude and tropics MLS day/night difference BrO data versus SLIMCAT. Data points represent seasonal 10° zonal means, as in Figure 9c, for pressures ranging from 10 to 3.2 hPa. Different symbols represent the four pressure surfaces (diamond, 10 hPa; triangle, 6.4 hPa; square, 4.6 hPa; cross, 3.2 hPa), and different colors represent the 11 latitude bins (red, 50°S ; orange, 40°S ; yellow, 30°S ; green, 20°S ; cyan, 10°S ; blue, 0° ; magenta, 10°N ; light gray, 20°N ; gray, 30°N ; dark gray, 40°N ; black, 50°N). The MLS error bars show the scatter ($\pm 1\sigma$ standard deviation) of the daily zonal means about the seasonal average for three representative latitudes and pressures.

expected to have been photochemically converted to inorganic bromine at altitudes above 25 km, even for cases where the balloon measurements do not overlap the MLS measurements, the Br_y comparison is expected to be meaningful. However, because stratospheric bromine loading has changed over time [Montzka *et al.*, 2003; WMO, 2003; Dorf *et al.*, 2006b], Br_y depends on the age of air, which varies with both altitude and latitude. Thus, to account for the age of air differences in the measurements, we also compare them on a plot of Br_y vs. N_2O .

[33] We infer Br_y vertical profiles from the measured BrO profiles using a Photochemical Steady State (PSS) box model. The PSS model is constrained to measurements of BrO, O_3 , and NO_2 , as described by Sioris *et al.* [2006] and Livesey *et al.* [2006a]. Since MLS does not measure NO_2 , we estimate it using MLS measurements of N_2O to specify NO_y (based upon well-established tracer relations) and PSS estimates of the NO_2/NO_y ratio. We use JPL 2002 kinetics [Sander *et al.*, 2003] with the addition of the reactions $\text{BrONO}_2 + \text{O} \rightarrow \text{BrO} + \text{NO}_3$ [Soller *et al.*, 2001], and $\text{BrO} + \text{OH} \rightarrow \text{Br} + \text{HO}_2$ [Bedjanian *et al.*, 2001], both of which are included in JPL 2006 [Sander *et al.*, 2006]. In addition, we update all reaction rate constants for reactions involving

bromine to JPL 06 values. Preliminary runs with JPL 2006 kinetics show that for the bromine family, model results using JPL 2002 kinetics with these changes are in agreement with those using JPL 2006 kinetics. The uncertainty in Br_y was determined by performing a sensitivity analysis with the PSS model, individually varying the abundances of BrO and NO_2 , as well as the rate constants for production and loss of BrO, and then adding the resulting changes in Br_y in quadrature [Sioris *et al.*, 2006].

[34] Vertical profiles of BrO obtained by the satellite-borne instrument SCIAMACHY (Scanning Imaging Absorption Spectrometer for Atmospheric Cartography) are available as well [Sinnhuber *et al.*, 2005; Sheode *et al.*, 2006; Sioris *et al.*, 2006]. However, the two sets of SCIAMACHY BrO data, obtained using two different retrieval methods, do not agree with each other [WMO, 2007, section 2.5.2.1]. In section 7.5 we present a comparison of Br_y inferred from MLS measurements of BrO to the two sets of published values of Br_y that have been inferred from SCIAMACHY BrO retrievals.

7.1. Comparison With DOAS BrO

[35] We first consider measurements of stratospheric BrO obtained by the remote-sensing, balloon-borne, Differential

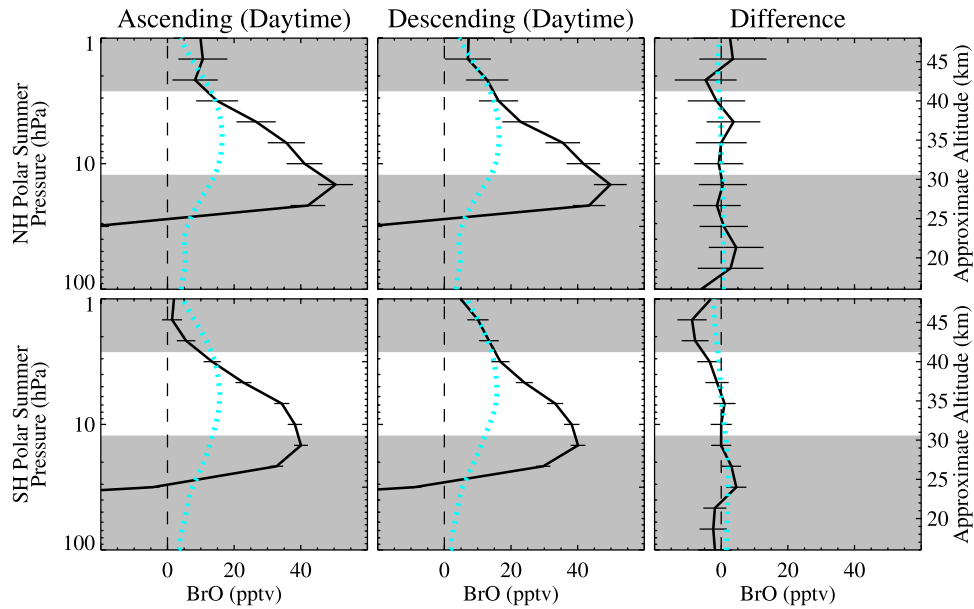


Figure 11. MLS v2.2 BrO vertical profiles for polar summer for ascending (daytime) measurements, descending (also daytime) measurements, and ascending/descending differences (which are expected to be zero). The profiles are seasonal averages of daily zonal means; the error bars show the precision. (top) NH mean for all days (4) reprocessed with MLS v2.2 in June, July, and August 2005 in the latitude range 75°N to 82°N and (bottom) SH zonal mean for all days (22) reprocessed with MLS v2.2 in December, January, and February 2004–2005 in the latitude range 75°S to 82°S. Regions shaded with dark gray indicate where MLS BrO has been determined to be unsuitable for scientific use. Also shown are corresponding SLIMCAT model profiles (cyan).

Optical Absorption Spectrometer (DOAS) instrument [Pfeilsticker *et al.*, 2000; Dorf *et al.*, 2006a]. This instrument detects ultraviolet (UV) and visible light from direct sun measurements during ascent of the balloon and solar occultation. Figure 14 (left) shows a vertical profile of DOAS measurements of BrO (green points) obtained on

23 March 2003 at 68°N [Dorf *et al.*, 2006a] during ascent of the balloon. This flight sampled air both inside and outside the vortex. The solar zenith angle (SZA) for these measurements changes with altitude, ranging from 82° to 88°. This version of the profile was obtained by smoothing the SCDs (Slant Column Densities) with a Gaussian of full-width-at-

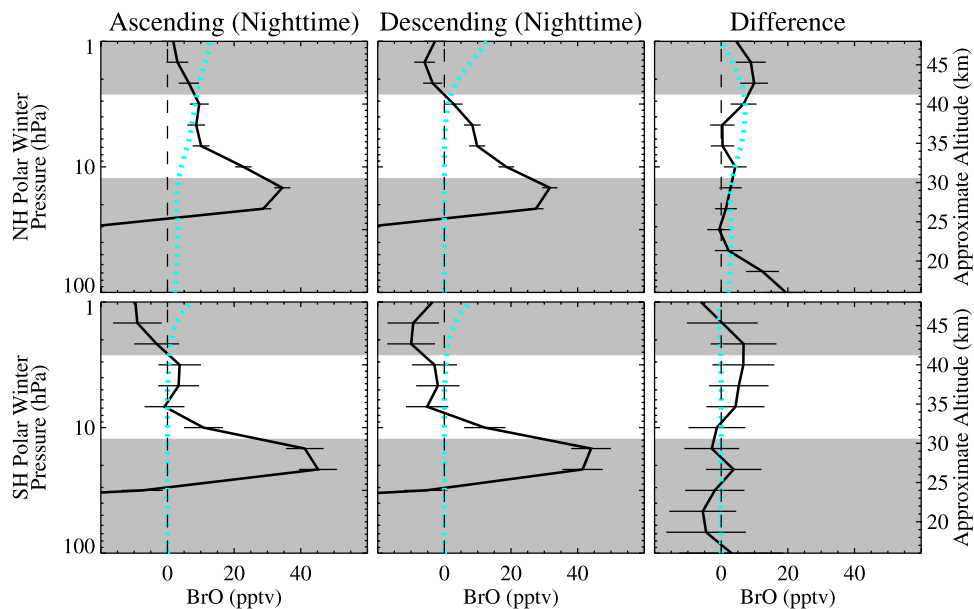


Figure 12. Same as Figure 11 except for polar winter. Regions shaded with dark gray indicate where MLS BrO has been determined to be unsuitable for scientific use.

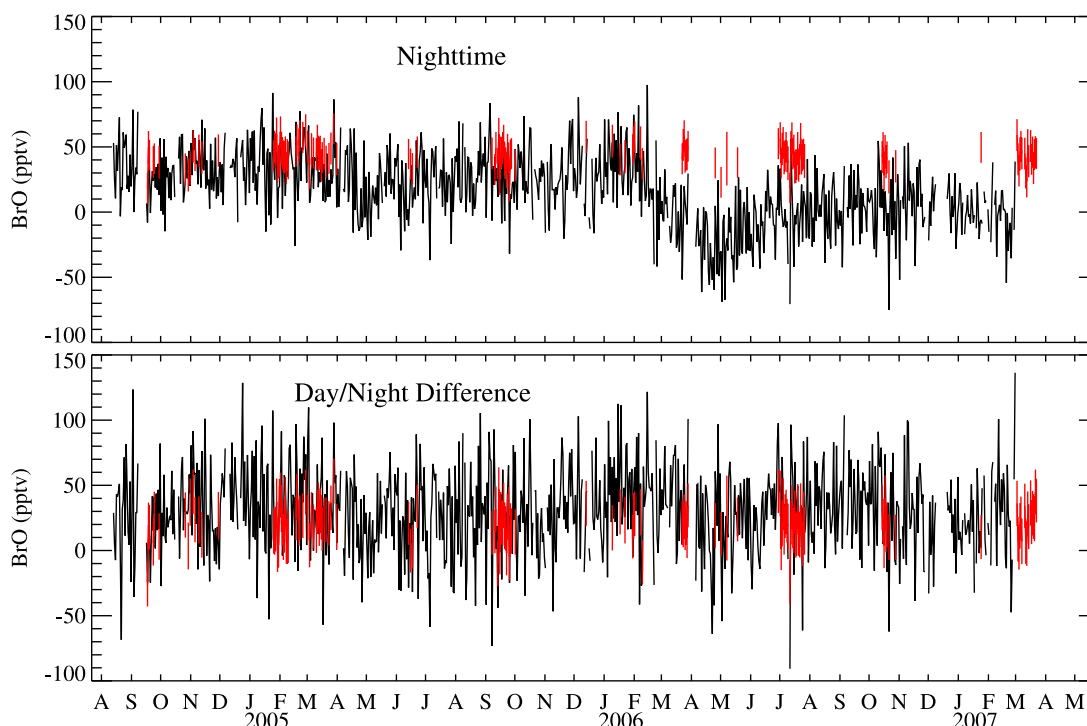


Figure 13. Time series of daily zonal mean BrO from 35°N to 45°N for the 10 hPa pressure surface averaged over the duration of the mission. Version 1.5 is shown in black, and v2.2 is shown in red. (top) Nighttime BrO measurements and (bottom) day/night differences. Error bars are shown for v2.2 but not for v1.5. The discontinuity in v1.5 nighttime data seen when band 13 was switched off in March 2006 is not seen in v2.2 nighttime data.

half-max (FWHM) of 1km before doing the profile inversion [Dorf, 2005]. The vertical profile of MLS BrO day/night difference is shown by the black points; the error bars indicate a combination of accuracy and precision. Only those MLS data points in the unshaded region should be compared with the DOAS measurements.

[36] We note some problems with comparing these DOAS measurements of BrO with the MLS measurements. First, the DOAS measurements were not obtained in the altitude region where the MLS BrO day/night difference measurements are suitable for scientific study (the unshaded region in Figure 14). Second, the high latitude (68°N) sampled by the DOAS instrument is not included in the average over the MLS measurements, which covers latitudes ranging from 55°S to 55°N. Third, as the two instruments take measurements at different solar zenith angles, we would like to use the PSS model to scale the DOAS BrO, measured at SZAs ranging from 82° to 88°, to the average SZA of the MLS measurements, about 40°. However, the smallest SZA attained at 68°N in March is 65°. Thus the dotted green line shows DOAS BrO scaled to a SZA of 65°, the closest possible to that of MLS BrO. These problems in nonoverlapping altitude range, different SZAs, and different latitudes are overcome by comparing the total inorganic bromine (Br_y), which is inferred from the measured BrO using the PSS model.

[37] Figure 14 (right) compares Br_y inferred by DOAS BrO (green dash), to that inferred by MLS BrO day/night difference (black dash). To infer DOAS Br_y , the PSS model

is constrained to DOAS measurements of NO_2 and O_3 . While there are no altitudes for which the DOAS measurements overlap the window of scientifically useful MLS measurements, at this high up in the stratosphere all the organic bromine should have been photochemically converted to Br_y , and so a comparison of Br_y is still relevant. We see good agreement between the two instruments for inferred Br_y , 22.1 ± 4 ppt for DOAS when averaged over altitude at and above 25 km, and 22.1 ± 5.5 ppt for MLS when averaged from 10 to 4.6 hPa. However, since Br_y depends on the age of air, which depends on both altitude and latitude, a more meaningful comparison is revealed by plotting the inferred Br_y vs. the tracer N_2O , which will be shown below.

7.2. Comparison With SAOZ-BrO BrO

[38] Figure 15 (left) shows a vertical profile of BrO (blue points) obtained by the balloon-borne, remote sensing SAOZ-BrO instrument during afternoon ascent of the balloon on 24 August 2004 at 52°N. SAOZ-BrO is also a solar occultation UV/vis spectrometer [Pundt et al., 2002]. For the SAOZ-BrO measurements, the solar zenith angle ranged from 80° to 90°. The dotted blue curve is the SAOZ-BrO values scaled to the average SZA (40°) of the MLS measurements using the PSS model. MLS BrO day/night difference measurements are shown by the black points, as in Figure 14. Within the window over which the scientifically useful MLS day/night difference measurements overlap the SAOZ measurements, the measurements agree within respective uncertainties.

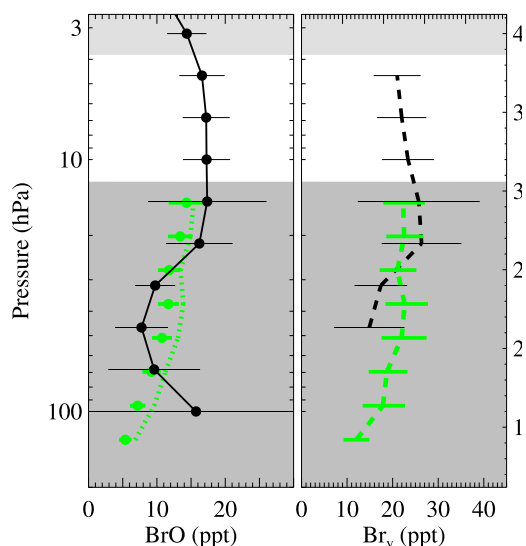


Figure 14. (left) Comparison of MLS day/night difference in BrO and (right) inferred Br_y with DOAS data from the 23 March 2003 flight at 68°N . In Figure 14 (left), the black points show the MLS BrO day/night difference averaged over all days (149) retrieved with v2.2 as of the writing of this manuscript, from 55°S to 55°N . Both the MLS BrO error bars and the DOAS error bars [Dorf et al., 2006a] indicate a combination of precision and accuracy. The average solar zenith angle of the MLS measurements is 40° . The nonshaded region indicates where the day/night difference in MLS BrO is representative of the daytime BrO. The lightly shaded region indicates where the day/night difference in MLS measurements of BrO is not representative of daytime BrO; rather, one needs to take account of the nonnegligible nighttime BrO, which we have not done here. The darker shaded region indicates where MLS measurements of BrO are not suitable for scientific use. The green points show DOAS BrO, while the green dotted curve shows that data scaled using a photochemical model to a SZA of 65° , the smallest SZA attained in March at that latitude. In Figure 14 (right), the dashed black curve is Br_y inferred from MLS BrO using a photochemical model. The dashed green curve shows Br_y similarly inferred from DOAS BrO. The error bars on the Br_y profiles indicate uncertainty due to measured BrO and NO_2 , as well as uncertainty in the model kinetics.

[39] Figure 15 (right) shows Br_y inferred by SAOZ-BrO BrO (blue dash), as well as that inferred by MLS BrO (black dash). To infer SAOZ Br_y , the PSS model is constrained to SAOZ measurements of NO_2 and O_3 . MLS-inferred Br_y agrees with SAOZ-inferred Br_y within the respective uncertainties. It should be kept in mind, however, that because of the difference in latitude, SAOZ at 52°N while MLS ranging from 55°N to 55°S , the average age of air will be quite different for the two profiles, and thus a more meaningful comparison will be shown below when Br_y is plotted against N_2O .

7.3. Comparison With SLS BrO

[40] Figure 16 (left) shows a vertical profile of BrO (red points) obtained by the balloon-borne Submillimeter Limb

Sounding (SLS-2) instrument on 20 September 2005 at 34°N . This is the first vertical profile of BrO obtained by this instrument, and thus is somewhat preliminary. SLS-2 is a cryogenic heterodyne instrument that detects atmospheric emission [Stachnik et al., 1992]. For BrO it detects the 624.768 GHz lines, corresponding to the ^{81}BrO isotope, the same lines detected by MLS, though MLS also uses the 650 GHz lines. The dotted red curve is the SLS-2 BrO scaled to the local time of the MLS measurements using the PSS model; in this case, since the SLS-2 measurements are obtained in the daytime (SZA = 51°), the scaling makes almost no change to the measured profile. In the window over which the useful MLS measurements overlap the SLS-2 measurements, the SLS BrO agrees with MLS at 10 hPa, but exceeds MLS BrO by 6 ppt at 4.6 hPa, which is outside the range of uncertainty.

[41] Figure 16 (right) shows Br_y inferred by SLS-2 BrO (red dashed) and that inferred by MLS BrO (black dashed). To infer SLS-2 Br_y , the PSS model is constrained to FIRS-2 measurements of NO_2 and O_3 . Comparing inferred Br_y , MLS and SLS-2 agree within uncertainty. As with the other comparisons, a more meaningful comparison will be shown below when Br_y is plotted against N_2O .

7.4. Comparison of Br_y in Tracer Space

[42] Figure 17 shows a plot of Br_y inferred from measurements of BrO by all four instruments in tracer space (i.e., plotted against N_2O). The advantage of this plot is that it accounts for differences in the age of air. The values of Br_y inferred from MLS BrO are plotted versus MLS measurements of N_2O . For the DOAS curve, the N_2O was obtained by the Limb Profile Monitor of the Atmosphere (LPMA) instrument, which flew on the same gondola as DOAS. For the SAOZ curve, the N_2O is a monthly zonal mean of MLS N_2O data for that month for the latitude bin 45°N to 55°N . For the SLS-2 curve, the N_2O was obtained by the FIRS-2

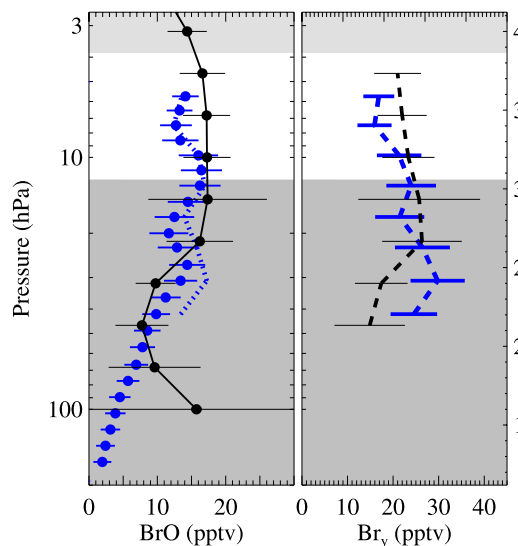


Figure 15. Same as Figure 14 except for BrO measurements obtained by the SAOZ-BrO instrument (blue) on 24 August 2004 at 52°N . The MLS BrO error bars indicate a combination of precision and accuracy, and the SAOZ error bars indicate precision.

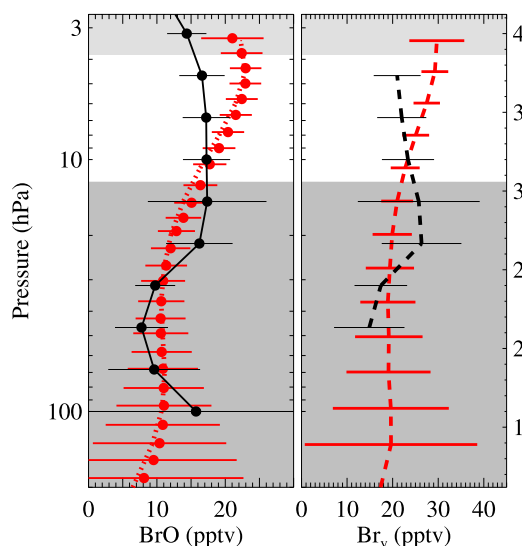


Figure 16. Same as Figure 14 except for BrO measurements obtained by the SLS-2 instrument (red) on 20 September 2005 at 35°N. The MLS BrO error bars indicate a combination of precision and accuracy, and the SLS-2 error bars indicate precision. Error bars on MLS and SLS-2 Br_y have been slightly offset for clarity.

instrument [Chance *et al.*, 1996] aboard the same gondola as SLS-2. Also shown is the “expected” relation (solid cyan) assuming production of Br_y from only the known organic precursors CH₃Br and halons [WMO, 2003]. To generate this curve we took the relation in the work by Wamsley *et al.* [1998] (dashed cyan), which was based on lower-stratospheric measurements of organic source gases obtained in 1994. We then updated the curve to the year of stratospheric entry relevant for this comparison, 2000 (solid line), by scaling the N₂O and Br_y to values appropriate for the year 2000 [WMO, 2007]. In the unshaded region of Figure 17, MLS-inferred Br_y is in agreement with all three of the other measurements within uncertainty. We also note that in this region, the Br_y inferred from all of the instruments exceeds that expected by the organic source gases. We discuss the implications of this excess Br_y in the next section.

7.5. Comparison With SCIAMACHY

[43] Although the MLS and SCIAMACHY data do not overlap in altitude (SCIAMACHY data is obtained in the lower stratosphere, from about 15 to 30 km) we compare the values of total bromine, Br_y, inferred from the measured BrO. SCIAMACHY retrievals by Shoede *et al.* [2006] infer 18.5 ± 4 ppt, consistent with an earlier estimate of Br_y from SCIAMACHY BrO by the same group [Sinnhuber *et al.*, 2005]. Retrievals of SCIAMACHY BrO by Sioris *et al.* [2006] infer 24.4 ± 2 ppt. The Br_y inferred from MLS BrO is 22.1 ± 5.5 ppt, falling between the two SCIAMACHY retrieval values.

8. Summary and Implications

[44] As summarized in Table 2, Aura MLS measurements of BrO using the v2.2 retrieval are scientifically useful for

pressures ranging from 10 hPa to 4.6 hPa, provided day/night differences are taken to remove biases, and significant averaging is done to reduce scatter. A maximum pressure of 10 hPa is chosen because large biases are seen in the nighttime MLS BrO signal at higher pressures. The minimum pressure of 4.6 hPa is chosen because that is the highest pressure for which day/night differences are a good measure of daytime BrO. While the useful pressure range can be extended to the 3.2 hPa pressure surface, the day/night difference will be a good measure of daytime BrO only if the nonnegligible nighttime BrO is taken into account, such as by using a photochemical model to compute the expected diurnal variation. At pressures lower than the 3.2 hPa surface, MLS BrO data are judged unsuitable for scientific study because the day/night differences in the summer polar regions do not behave as expected. For a monthly 10° zonal mean, over the pressure range from 10 to 3.2 hPa, the precision in the day/night difference is ± 4 ppt and the uncertainty in the accuracy is about $\pm 20\%$. The vertical resolution between 10 and 3.2 hPa is about 5.5 km.

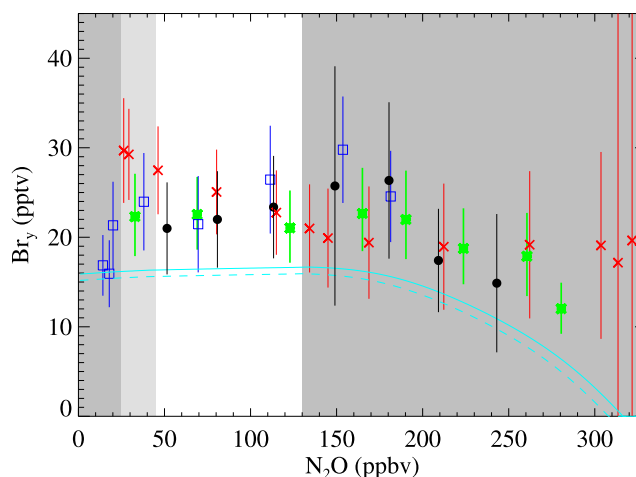


Figure 17. Tracer space comparison of Br_y inferred from measurements of BrO obtained by MLS and three balloon instruments. The MLS Br_y (black) is inferred from daily zonal mean BrO using a photochemical model and then averaged over latitudes ranging from 55°S to 55°N and over all 149 v2.2 days processed to date. The DOAS (green) Br_y is inferred from BrO measured during the March 2003 Kiruna flight, the SAOZ-BrO Br_y (blue) is inferred from BrO measured during the August 2004 flight, and the SLS-2 inferred Br_y (red) is for the September 2005 flight. Also shown is the expected relation assuming known organic precursors, based on measurements in 1994 [Wamsley *et al.*, 1998] (dashed cyan) and updated to the year of stratospheric entry of 2000 (solid cyan) [WMO, 2007]. The nonshaded region indicates where the day/night difference in MLS BrO is representative of the daytime BrO. The lightly shaded region, as well as lower N₂O, indicates where the day/night difference in MLS measurements of BrO is not representative of daytime BrO; rather, one needs to take account of the nonnegligible nighttime BrO. The darker shaded regions indicate where MLS measurements of BrO are not suitable for scientific use.

[45] To compare our current retrieval, v2.2, with the interim retrieval used in a previous study [Livesey *et al.*, 2006a], we repeat their calculation of total inorganic bromine, Br_y. In that study, Br_y was calculated from the MLS interim-version BrO using the PSS model, and an annual average was taken over latitudes ranging from 55°S to 55°N and pressures ranging from 10 hPa to 2.2 hPa. While day/night differences were used over the pressure range 10 hPa to 4.6 hPa, the ascending measurements were used for the 3.2 and 2.2 hPa pressure levels. Here we again use the PSS model to calculate Br_y from MLS BrO, except we use v2.2 BrO. We take an average over the same latitude range but only over the days (149) retrieved with v2.2 at the time of writing this manuscript. We use only day/night differences, and average over a smaller pressure range, from 10 hPa to 4.6 hPa, where the day/night difference is a good measure of daytime BrO. We find a total stratospheric bromine, Br_y, of 22.1 ± 5.5 ppt, which is in acceptable agreement with the Livesey *et al.* value of 18.6 ± 5.5 ppt, within the respective uncertainties of each value.

[46] As seen in Figure 17, the Br_y inferred from MLS measurements of BrO exceeds the expected value based on the organic precursors CH₃Br and halons. This is in agreement with mounting evidence that very short lived source gases (VSLs) are contributing significantly to the bromine budget [Pfeilsticker *et al.*, 2000; Salawitch *et al.*, 2005; Dorf *et al.*, 2006b; WMO, 2007]. If we use an estimate of 15.6 ppt for the total bromine abundance of the source gases CH₃Br and halons present in the atmosphere at the time of the MLS measurements [Montzka *et al.*, 2003; Livesey *et al.*, 2006a], we estimate that the VSLs contribute about 6.5 ± 5.5 ppt to the bromine budget. Moreover, the agreement of MLS day/night difference BrO with the SLIMCAT model, seen in Figure 10, also supports this interpretation, since the SLIMCAT model assumed a total inorganic bromine of 22 ppt, which represents a contribution of 6 ppt from VSLs.

[47] Our plans for the future include improving the lower-altitude MLS retrievals, where the vertical distribution of stratospheric bromine is currently a controversial issue. While we have some understanding as to the sources of large bias in the MLS BrO data for pressures of 15 hPa and greater, we plan to work on reducing the effect of these sources. In particular, we are looking into a new version of the radiance-based retrieval (called “interim version” in this manuscript) that will be optimized for BrO. (In contrast, the retrieval v2.2 is designed to optimize for the whole of the 640 GHz radiometer.) We also plan to examine retrievals of BrO in the polar regions, where for summer and winter seasons we cannot use day/night differences. It may be possible to characterize the bias as a function of altitude both for midlatitudes as well as for the polar regions in the spring, and then use this result to compensate for the bias in the polar regions during summer and winter, as was done for MLS ClO by M. L. Santee *et al.* (Validation of the Aura Microwave Limb Sounder ClO measurements, submitted to *Journal of Geophysical Research*, 2007). We do not understand why the day/night difference for the southern polar summer region did not go to zero for pressures of 2.2 and less. We plan to further investigate this issue as well.

[48] **Acknowledgments.** Research at the Jet Propulsion Laboratory, California Institute of Technology, is performed under contract with NASA. We greatly appreciate the efforts of Bojan Bojkov and the Aura Validation Data Center (AVDC) team, whose work facilitated the MLS validation activities. Thanks to the Aura Project for their support throughout the years (before and after Aura launch), in particular M. Schoeberl, A. Douglass (also as cochair of the Aura validation working group), E. Hilsenrath, and J. Joiner. We also acknowledge the support from NASA Headquarters, P. DeCola for MLS and Aura. The SAOZ team thanks M. van Roozendaal (Belgian Institute for Space Aeronomy) for spectra processing, the Canadian Space Agency (CSA), Environment Canada, Scientific Instrumentation Limited, the MANTRA Pls K. Strong and S. Melo (University of Toronto) for flying SAOZ during the MANTRA 2004 campaign, and Centre National d’Etudes Spatiales (CNES) for its financial support.

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